


Last time, we got an introduction to the partition function  $Z$

$$(1.1) \quad Z = \sum_s e^{-\epsilon_s / \tau}$$

$\epsilon_s$  ← energy of microstate labelled by  $\{s\}$   
 $s$  ← MICRO STATES

Then

$$(1.2) \quad P_s = \frac{e^{-\epsilon_s / \tau}}{Z}$$

Today, we'll learn about more ways in which  $Z$  can be very, VERY useful. 

- Consider the average energy of our system  $S$  in thermal contact, with reservoir  $R$  at temperature  $\tau$ .  
 ↙ exchange of energy only (not particles which make up our system  $S$  of interest)

We know the abstract formula for the thermal average of any  $X$ :

$$(1.3) \quad \langle X \rangle = \sum_s X_s \cdot \left( \frac{e^{-\epsilon_s / \tau}}{Z} \right)$$

So for energy we have

$$(1.4) \quad \langle \epsilon \rangle = \sum_s \epsilon_s \frac{e^{-\epsilon_s / \tau}}{Z}$$

$\langle \epsilon \rangle$  is traditionally called  $U$ . ← "big  $U$ "

$$(1.5) \quad \Rightarrow \quad U = \sum_s \epsilon_s \frac{e^{-\epsilon_s / \tau}}{Z}$$

Now let's do some mathematical manipulations, with the physics goal of finding the average energy.

②

We know  $U = \sum_s \epsilon_s \left( \frac{e^{-\epsilon_s/\tau}}{Z} \right)$   
 $= \frac{1}{Z} \sum_s \epsilon_s e^{-\epsilon_s/\tau}$  ( $\because Z$  independent of  $s$ )

Now what?!

Well, the goal is to use  $Z$  to find  $U$ .  
What operations do we know how to do to  $Z$ ?

▷ We can take derivative(s), for one example...


But we should take the derivative w.r.t a macroscopic variable

So try  $\tau$ :  $Z$  depends on this explicitly.

(2.1)  $\frac{\partial Z}{\partial \tau} = \frac{\partial}{\partial \tau} \sum_s e^{-\epsilon_s/\tau}$   
 $= \sum_s \frac{\partial}{\partial \tau} e^{-\epsilon_s/\tau}$   
 (2.2)  $= \sum_s \left( + \frac{\epsilon_s}{\tau^2} \right) e^{-\epsilon_s/\tau}$   
 (2.3)  $= \frac{1}{\tau^2} \sum_s \epsilon_s e^{-\epsilon_s/\tau}$   
 aha! 😊

So... we've discovered that

(2.4)  $\frac{\partial Z}{\partial \tau} = \frac{1}{\tau^2} (U Z)$

(2.5) i.e.   $U = \frac{\tau^2}{Z} \frac{\partial Z}{\partial \tau}$

↖ This works for completely general system  $\mathcal{S}$  in thermal contact with  $\mathcal{R}$  @ temperature  $\tau$  😊

## Example

As we found last time,

$$(3.1) \quad Z = 1 + 3e^{-\epsilon/\tau}$$

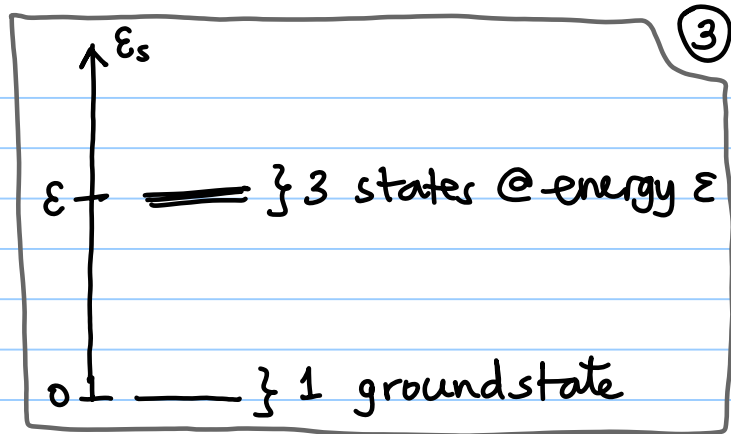
$$\text{So } \frac{\partial Z}{\partial \tau} = 3e^{-\epsilon/\tau} \cdot \left(\frac{+\epsilon}{\tau^2}\right)$$

$$(3.2) \quad = \frac{3\epsilon}{\tau^2} e^{-\epsilon/\tau}$$

$$\text{hence } U = \frac{\tau^2}{Z} \frac{\partial Z}{\partial \tau}$$

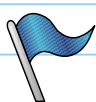
$$= \frac{\tau^2}{(1+3e^{-\epsilon/\tau})} \cdot \frac{3\epsilon}{\tau^2} e^{-\epsilon/\tau}$$

$$(3.3) \quad \text{re. } U = \frac{3\epsilon e^{-\epsilon/\tau}}{(1+3e^{-\epsilon/\tau})}$$
$$= \frac{e^{-0} \cdot 0}{(1+3e^{-\epsilon/\tau})} + \frac{3e^{-\epsilon/\tau} \cdot \epsilon}{(1+3e^{-\epsilon/\tau})}$$
$$= 0 P_0 + 3 P_\epsilon \quad \checkmark$$



## Heat Capacity

(3.4)



$$C_V = \left(\frac{\partial U}{\partial \tau}\right)_V$$

← how much energy increase you get if you crank up  $\tau$  😊  
← dimensionless

Here, for our example,

$$C_V = \frac{\partial}{\partial \tau} \left\{ \frac{3\epsilon e^{-\epsilon/\tau}}{(1+3e^{-\epsilon/\tau})} \right\}$$

$$\Rightarrow C_V = \frac{[(1+3e^{-\epsilon/\tau}) \cdot (\epsilon/\tau^2)e^{-\epsilon/\tau}] - [e^{-\epsilon/\tau} \cdot 3(\epsilon/\tau^2)e^{-\epsilon/\tau}]}{3\epsilon (1+3e^{-\epsilon/\tau})^2}$$

$$= \frac{1}{(1+3e^{-\epsilon/\tau})^2} \frac{\epsilon}{\tau^2} e^{-\epsilon/\tau} ([1 + \cancel{3e^{-\epsilon/\tau}}] - [3\cancel{e^{-\epsilon/\tau}}])$$

$$(3.5) \quad C_V = \frac{3\epsilon^2 e^{-\epsilon/\tau}}{\tau^2 (1+3e^{-\epsilon/\tau})^2}$$

Notice that again, this is a function of  $(\epsilon/\tau)$  only.

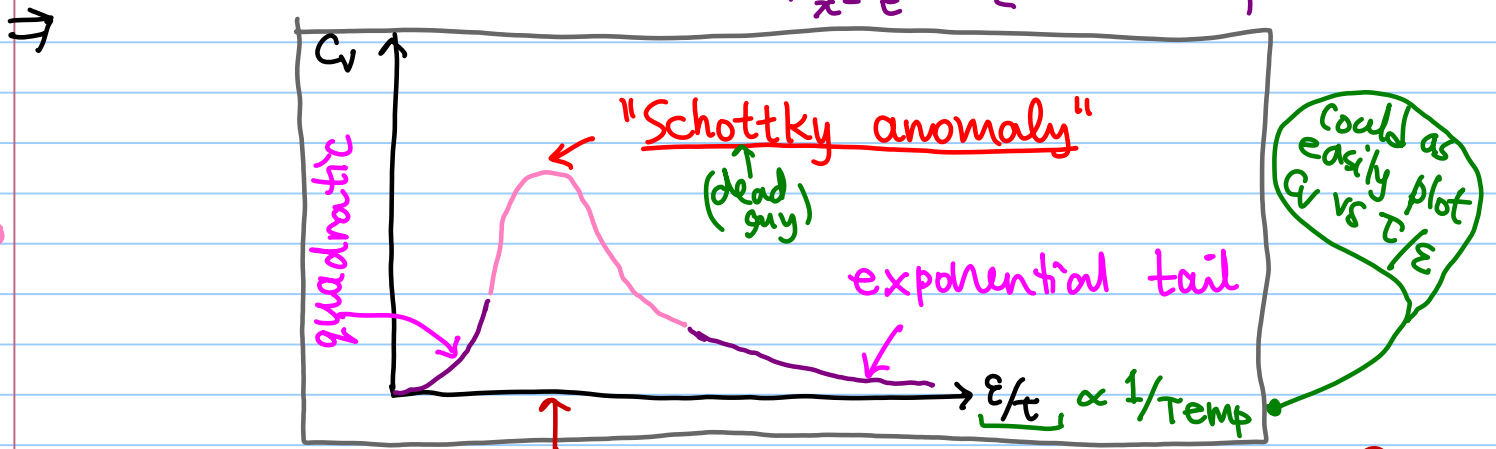
Let's plot it.

• First, consider  $\frac{\epsilon}{T} \rightarrow 0$ .

(4.1) Then  $C_V \rightarrow 3 \left(\frac{\epsilon}{T}\right)^2 \cdot \frac{1}{(1+0)^2} = 3 \left(\frac{\epsilon}{T}\right)^2 \rightarrow 0$  (quadratically)

Second, consider  $\frac{\epsilon}{T} \rightarrow \infty$ .

(4.2) Then  $C_V \rightarrow 3 \left(\frac{\epsilon}{T}\right)^2 \frac{e^{-\epsilon/T}}{(1+0)^2} = 3 \left(\frac{\epsilon}{T}\right)^2 e^{-\epsilon/T} \rightarrow 0$  (exponential always)   
 "x<sup>2</sup> e<sup>-x</sup>" ← (beats power-law!)



biggest energy bang for your temperature buck 😊

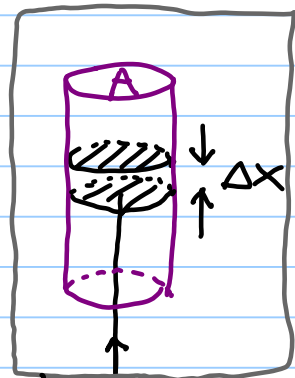
Work and Pressure

← We all know a bit about these, eh? 😊

Suppose we had a piston.

Let it have cross-sectional area A, and suppose it moves a distance Δx.

(4.4) We know  $dW = \vec{F} \cdot d\vec{x}$   
(4.5)  $\Rightarrow \Delta W = p \cdot A \cdot \Delta x$   
 $= p \Delta V$



This formula generalizes to 3-d expansions too

(5.1) Generally,  $\Delta U = \Delta Q - \Delta W$   
so  $\Delta U = 0 - p\Delta V$  if  $\sigma = \text{constant}$

(5.2)  $\Rightarrow p = -\left(\frac{\partial U}{\partial V}\right)_\sigma$

If  $U$  &  $V$  were the two independent variables for our thermodynamic system,  $\sigma = \sigma(U, V)$ .  
(Generally, keeping  $\sigma = \text{const.}$  is experimentally tricky!)

(5.3)  $d\sigma = 0$   
says  $\left(\frac{\partial \sigma}{\partial U}\right)_V dU + \left(\frac{\partial \sigma}{\partial V}\right)_U dV = 0$

So \* at constant  $\sigma$  only \* can "divide by  $dV$ "  
 $\Rightarrow \left(\frac{\partial \sigma}{\partial U}\right)_V \left(\frac{\partial U}{\partial V}\right)_\sigma + \left(\frac{\partial \sigma}{\partial V}\right)_U = 0$   
i.e.  $\frac{1}{\tau} \cdot -p + \left(\frac{\partial \sigma}{\partial V}\right)_U = 0$

(5.4) Hence  $p = \tau \left(\frac{\partial \sigma}{\partial V}\right)_U$

← These are equivalent

In general,  $d\sigma = \left(\frac{\partial \sigma}{\partial U}\right)_V dU + \left(\frac{\partial \sigma}{\partial V}\right)_U dV = \frac{1}{\tau} dU + \frac{p}{\tau} dV$

(5.5) i.e.  $dU = \tau d\sigma - p dV$  First Law

### Free Energy

Suppose magician destroyed rabbit(!)



What happens to energy?

(5.6)

← [Thanks to Dan Schroeder for pic.]

We expect that it won't be possible to capture the whole energy and put it to work - we will lose some heat to the surroundings (which effectively constitute a thermal reservoir).

$$\Delta(\text{energy}) \rightarrow (\text{work done}) + (\text{heat lost})$$

Traditionally,  $\begin{cases} U = \text{energy of system} \\ Q = \text{heat added to system} \\ W = \text{work done by system} \end{cases}$

(6.1)  $\rightarrow \therefore \Delta U = \Delta Q - \Delta W$   
 (6.2) or  $\boxed{dU = dQ - dW}$

▷ What is heat? It is macroscopic, a form of energy.

(6.3)  $\boxed{dQ = \tau d\sigma} = T dS$  ← "tedious" (haha!) <sup>Pron.</sup>

so, for our rabbit-destruction process,  $\Delta Q = \tau \Delta \sigma$  since R keeps  $\tau = \text{const.}$

$\therefore 0 - Q = \tau(\sigma - 0)$  i.e.  $\Delta Q = -\tau\sigma$

$\Rightarrow 0 - U = -\tau\sigma - \Delta W$   
 or  $\Delta W_{\text{rabbit}} = U - \tau\sigma$

This is the energy available, or "free", for doing work. Cool! This is what we wanted. 😊

(6.4) Helmholtz free energy  $\boxed{F = U - \tau\sigma}$

An equivalent\* definition, which I <sup>strongly</sup> prefer, is

$\boxed{F = -\tau \log Z}$  \*

• Since  $dF = dU - \tau d\sigma - \sigma d\tau = -pdV - \sigma d\tau$   $(F = U - \tau\sigma)$   
 $(dU = \tau d\sigma - pdV)$

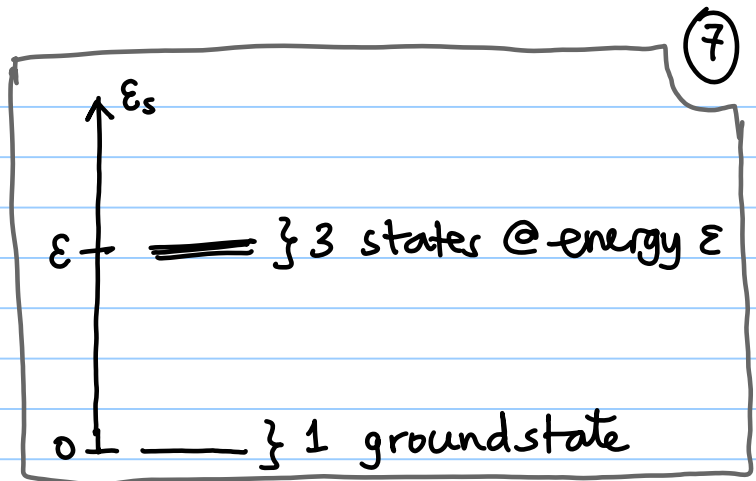
$\Rightarrow \boxed{P = -\left(\frac{\partial F}{\partial V}\right)_{\tau}} *$   $\boxed{\sigma = \left(\frac{\partial F}{\partial \tau}\right)_{V}} *$

mega-useful!

\* see KK!

## Example?

For our system with  
we had  
 $Z = 1 + 3e^{-\epsilon/\tau}$ .



So

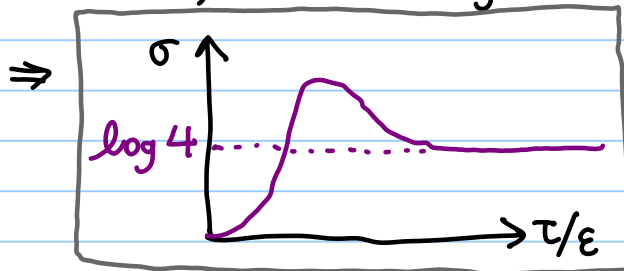
- $F = -\tau \log Z$

i.e.  $F = -\tau \log (1 + 3e^{-\epsilon/\tau})$

- $\left(\frac{\partial F}{\partial \tau}\right)_V = -\log(1 + 3e^{-\epsilon/\tau}) - \frac{\tau}{(1 + 3e^{-\epsilon/\tau})} \cdot \frac{+3\epsilon e^{-\epsilon/\tau}}{\tau^2}$   
 $\therefore \sigma = \log(1 + 3e^{-\epsilon/\tau}) + \frac{3(\epsilon/\tau)e^{-\epsilon/\tau}}{(1 + 3e^{-\epsilon/\tau})}$

As  $\underbrace{\epsilon/\tau}_{:=x} \rightarrow \infty$  ( $\tau \rightarrow 0$ ),  $\sigma \rightarrow 0 + \mathcal{O}(e^{-x})$

As  $\epsilon/\tau \rightarrow 0$  ( $\tau \rightarrow \infty$ ),  $\sigma \rightarrow \log(4) + \mathcal{O}(x)$



- $\left(\frac{\partial F}{\partial V}\right)_\tau = ?$  Well,  $Z = Z(\epsilon/\tau) = Z(\epsilon)$  @  $\tau = \text{const.}$

$\Rightarrow$  to find pressure, we need to know how  $\epsilon$  depends on  $V$ .

$\rightarrow$  We'll start that problem in the second half of next lecture 😊  
i.e. finding  $\left(\frac{\partial \epsilon}{\partial V}\right)_\tau$  - for ideal gas.