

Light-Front Coordinates

①

For purposes of quantizing our string (later on), we will need to get familiar with a particular choice of (t, \mathbf{x}) different from the static gauge studied so far.

Here is a lightning ⚡ review of these concepts.

$$(1.1) \quad \text{Define } x^\pm \equiv \frac{1}{\sqrt{2}} (x^0 \pm x^1)$$

$$(1.2) \quad \text{Then } -(dx^0)^2 + (dx^1)^2 = -2dx^+dx^-$$

So that the spacetime metric is $\begin{pmatrix} 0 & -1 & 0 & 0 \\ -1 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} = (\eta_{\mu\nu})$

Light-cone components of any vector or tensor can easily be found. Also,

$$(1.3) \quad a \cdot b = -a^+ b^- - a^- b^+ + a^2 b^2 + a^3 b^3$$

In Light-cone coords, let us inspect what physicists call the mass shell condition (it's one of two invariants of Poincaré: $P^\mu P_\mu$.)

$$P^\mu P_\mu = \eta_{\mu\nu} P^\mu P^\nu = -P^- P^+ - P^+ P^- + \vec{P}_\perp^2$$

$$= -2P^- P^+ + \vec{P}_\perp^2 = -m^2$$

$$\Rightarrow P^- P^+ = \frac{1}{2} (m^2 + \vec{P}_\perp^2)$$

$$(1.4) \quad \boxed{-P_+ = P^- = \frac{1}{2P^+} (m^2 + \vec{P}_\perp^2)}$$

"looks" non-relativistic (!)

$$(1.5) \quad (\text{where } p^\pm = \frac{1}{\sqrt{2}} \left(\frac{E}{c}, \pm p^1, 0, 0 \right))$$

- Light-cone time is taken to be x^+ .
 $\Rightarrow p_+$ is the light-cone energy

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String gauge choices: noncovariant gauges

Linear combination of position fields set to $\lambda\tau$:-

$$(2.1) \quad \boxed{\eta_\mu X^\mu(\tau, \sigma) = \lambda\tau} \quad \text{Breaks Lorentz invariance (via choice of particular } \eta_\mu \text{).}$$

The set of points x^μ satisfying $\eta_\mu x^\mu = \lambda\tau$ forms a hyperplane normal to η_μ .
 \Rightarrow the curve where the string worldsheet intersects this hyperplane normal to η_μ is the string at fixed τ .

What does this imply about energy-momentum? Static gauge!
 The momentum density was $P^{\tau\mu}$, where energy is taken as " $\partial/\partial(\text{time})$ ". If instead we use the gauge (2.1); what we should demand is conservation not of $P^{\tau\mu}$ but

$$(2.2) \quad \eta_\mu P^{\tau\mu} = \text{constant in } \tau \text{ \& } \sigma$$

This makes the string have constant energy density along its length (tension is constant in σ parametrization); the eqn (2.2) also ensures τ and σ are nicely orthogonal as we demanded earlier. Zwiebach shows that for an open string we can take $\sigma \in [0, \pi]$.

Equation of motion?

$$\frac{\partial}{\partial\tau} P^{\tau\mu} + \frac{\partial}{\partial\sigma} P^{\sigma\mu} = 0$$

$$\Rightarrow \frac{\partial}{\partial\tau} (n \cdot P^\tau) = - \frac{\partial}{\partial\sigma} (n \cdot P^\sigma)$$

$$(2.3) \quad = 0 \quad \Rightarrow (n \cdot P^\sigma) \text{ independent of } \sigma.$$

For an open string, we required for free endpoints $P^{\sigma\mu} = 0$ at the endpoints. But this implies $(n \cdot P^\sigma) = 0$ at endpoints. Independence of σ then gives

$$(c) \quad n \cdot P^\sigma = 0 \text{ everywhere for open string (!)}$$

③

But for closed strings we have different BC's. Had Momentum. (conserved): $p_\mu = \int_\gamma (\rho_\mu^\tau d\sigma - \rho_\mu^\sigma d\tau)$

where γ winds once around cylindrical-topology worldsheet.

Now, choose purely spatial path along $\partial/\partial\sigma$:

(3-1) Let
$$(h \cdot p)^\sigma = \frac{2\pi}{\beta} \int_0^\sigma d\tilde{\sigma} \ n \cdot \rho^\tau(\tau, \tilde{\sigma})$$

\leftarrow to make $\sigma \in [0, 2\pi]$ for closed strings
 $\leftarrow \sigma \in [0, \pi]$ for open strings

while

(3-2)
$$n \cdot X(\tau, \sigma) = \beta \alpha' (h \cdot p) \tau$$

\leftarrow gauge choice with right dimensions, and tension, etc.

For the closed string, is $n \cdot p^\sigma = 0$ like for open string?

$$n \cdot p^\sigma = h_\mu \left\{ \frac{-1}{2\pi\alpha'} \left[\frac{(\dot{X} \cdot X') \partial_\tau X^\mu - \dot{X}^2 \partial_\sigma X^\mu}{\sqrt{(\dot{X} \cdot X')^2 - \dot{X}^2 (X')^2}} \right] \right\}$$
$$= \frac{-1}{2\pi\alpha'} \frac{(\dot{X} \cdot X') \partial_\tau (h \cdot X)}{\sqrt{(\dot{X} \cdot X')^2 - \dot{X}^2 (X')^2}} \quad \text{using (3-2)}$$

In order to make this vanish, we would need

(3-3)
$$\dot{X} \cdot X' = 0$$

on at least one point along the string; that would be enough to make it zero everywhere along σ by (2-3).

Note that (3-3) is a relativistic dot product.

Is it zero somewhere?

Let $\{t^M\}$ the tangent vector to the worldsheet at a point P then $\{x'^M\}$ and $\{t^M\}$ generate T_P ; we know they are not parallel because X'^M is spacelike while t^M is timelike.

If they are orthogonal, i.e. $t^M x'_\mu = 0$, then t^M is the tangent vector doing the job. Otherwise, define

$$v^M = t^M - \frac{(t \cdot X')}{X' \cdot X'} X'^M \quad \text{which is } \perp X'_\mu : v^M x'_\mu = 0.$$

Defining $\sigma=0$ to be the line given by

$$X^M(P) + \epsilon v^M$$

ensures that the tangent $\perp X'_\mu$ and, since the tangent vector $\propto \dot{X}^M$, have (3-3) satisfied. \Rightarrow

$$\Rightarrow \boxed{h \cdot p^\sigma = 0}$$

(open & closed)

σ can be any point on string.

$\frac{\partial}{\partial\sigma}$ symmetry, unfixed.

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(4.1) Having $\dot{x} \cdot x' = 0$ simplifies life.

$$\Rightarrow p^{\tau\mu} = \frac{1}{2\pi\alpha'} \frac{(x')^2 \dot{x}^\mu}{\sqrt{-(\dot{x})^2 (x')^2}} \Rightarrow n \cdot p^\tau = \frac{1}{2\pi\alpha'} \frac{(x')^2 (n \cdot \dot{x})}{\sqrt{-(\dot{x})^2 (x')^2}}$$

but $n \cdot x = \beta \alpha' (n \cdot p) \tau$ so $(n \cdot \dot{x}) = \beta \alpha' (n \cdot p)$

Also, since $(n \cdot p) \sigma = \frac{2\pi}{\beta} \int_0^\sigma d\tilde{\sigma} [p^\tau(\tau, \tilde{\sigma}) \cdot n]$

$$(n \cdot p) = \frac{2\pi}{\beta} n \cdot p^\tau \quad \text{so}$$

$$n \cdot p^\tau = \frac{\beta}{2\pi} (n \cdot p) = \frac{n \cdot \dot{x}}{2\pi\alpha'} = \frac{1}{2\pi\alpha'} \frac{(x')^2 n \cdot \dot{x}}{\sqrt{-(\dot{x})^2 (x')^2}}$$

(4.2) so $\Rightarrow \frac{(x')^2}{\sqrt{-(\dot{x})^2 (x')^2}} = 1 \Rightarrow \boxed{\dot{x}^2 + (x')^2 = 0}$

Combining (4.1) & (4.2) gives $\boxed{\dot{x} \pm x' = 0}$

Momenta simplify:

(4.3) $\boxed{p^{\tau\mu} = \frac{1}{2\pi\alpha'} \dot{x}^\mu}$

and

(4.4) $\boxed{p^{\sigma\mu} = -\frac{1}{2\pi\alpha'} x'^\mu}$

(4.5) and so $\boxed{\ddot{x}^\mu - (x^\mu)'' = 0}$ equation of motion

The solution to the e-o-m (4.5) can be Fourier-expanded with the centre of mass motion split off:

(4.6a) $\boxed{x^\mu(\tau, \sigma) = x_0^\mu + \sqrt{2\alpha'} \alpha_0^\mu \tau + i\sqrt{2\alpha'} \sum_{n \neq 0} \frac{1}{n} \alpha_n^\mu e^{-in\tau} \cos(n\sigma)}$

standing waves satisfy Neumann BC's @ endpoints.

where

(4.6b) $\boxed{\alpha_0^\mu \equiv \sqrt{2\alpha'} p^\mu}$

so that $(\dot{x}^\mu \pm x'^\mu) = \sqrt{2\alpha'} \sum_{n \in \mathbb{Z}} \alpha_n^\mu e^{-in(\tau \pm \sigma)}$

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Light-cone gauge (specific case)

We set $\eta_\mu = (\frac{1}{\sqrt{2}}, \frac{1}{\sqrt{2}}, 0, \dots, 0)$

so that $h \cdot X = X^+$ so $h \cdot p = p^+$ and

$$(5.1) \quad \boxed{X^+(\tau, \sigma) = \beta \alpha' p^+ \tau}, \quad \boxed{p^+ \sigma = \frac{2\pi}{\beta} \int_0^\sigma d\tilde{\sigma} P^{+\tau}(\tau, \tilde{\sigma})}$$

Then requiring $(\dot{X}^\pm \pm X'^i)^2 = 0$

$$(5.2) \quad \Rightarrow \quad \boxed{(\dot{X}^- \pm X'^i) = \frac{1}{\beta \alpha'} \frac{1}{2p^+} (\dot{X}^i \pm X'^i)^2}$$

Using the general mode expansion (4.6a) and (5.1), see that

$$(5.3) \quad \text{i.e. } \underline{\text{in light-cone gauge, } X^+ \text{ does not oscillate.}}$$

$x_0^+ = 0, \quad \alpha_n^+ = \alpha_n^- = 0; \quad p^+ = \sqrt{2\alpha'} \alpha_0^+$

Since we have (5.2), we can see that the Θ oscillators are going to be expressed in terms of quadratic combos of X^i oscillators.

$$(5.4) \quad \text{In detail, (see Zwiebach p.162), obtain [algebra] } \quad \boxed{\sqrt{2\alpha'} \alpha_n^- = \frac{1}{2p^+} \sum_{p \in \mathbb{Z}} \alpha_{n-p}^I \alpha_p^I} \quad (I \text{ is transverse})$$

So the α_n^- are not independent physical wiggles. It is the X^I that have the "real" dynamics. 😊

▷ Notice, for QFT students: when we set the gauge, (5.1), we got rid of two longitudinal degrees of freedom: not only did we lose X^+ oscillators via (5.1), the constraints knocked out X^- oscillators too. Quite generally, see that to knock out one longitudinal component of a vector, need to knock out a second. i.e. for A_μ in E&M also get only (D-2) indep. d.o.f. (polarizations)

Virasoro modes

(6.1) Rename $L_n^\perp \equiv \frac{1}{2} \sum_{p \in \mathbb{Z}} \alpha_{n-p}^I \cdot \alpha_p^I$

(6.2) then $2p^+ p^- = \frac{1}{\alpha'} L_0^\perp$

(6.3) so that $(\dot{X}^\pm \pm X'^\pm) = \frac{1}{p^+} \sum_{n \in \mathbb{Z}} L_n^\perp e^{-in(\tau \pm \sigma)}$

We can also easily compute the mass in light-cone gauge:

$$\begin{aligned} M^2 &= -p^\mu p_\mu = +p^+ p^- + p^- p^+ - p^I p^I \\ &= \frac{1}{\alpha'} L_0^\perp - p^I p^I \\ &= \frac{1}{\alpha'} \left[\frac{1}{2} \alpha_0^I \alpha_0^I + \sum_{n \in \mathbb{Z}^+} \alpha_n^{I*} \alpha_n^I \right] - p^I p^I \end{aligned}$$

(6.4) $M^2 = \frac{1}{\alpha'} \sum_{n \in \mathbb{Z}^+} n \alpha_n^{I*} \alpha_n^I \geq 0$ classically

where $\alpha_n^\mu = \alpha_n^\mu \sqrt{n}$, $n \geq 1$.

Later on, we will quantize the string.

Each mode will contribute a zero-point energy, and we have to figure out the effect on the $(\text{mass})^2$ - once we know how to quantize!

First, we discuss the simpler cases of scalar, vector and tensor fields in light-cone gauge to get some warm-up practice with the canonical quantization procedure.

Then we will graduate to quantizing strings.

Spin-zero

The relativistic action for a scalar, which is the minimal modification of "L=T-V" consistent with Lorentz invariance, is

$$(7.1) \quad S = \int d^D x \left\{ -\frac{1}{2} \partial^\mu \phi \partial_\mu \phi - \frac{1}{2} m^2 \phi^2 \right\}$$

This has canonical momenta so that

$$\Pi_\mu = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = -\partial_\mu \phi$$

$$\Pi_0 = \partial_0 \phi \quad \text{and so}$$

$$\mathcal{H} = \frac{1}{2} \Pi^2 + \frac{1}{2} (\nabla \phi)^2 + \frac{1}{2} m^2 \phi^2 \quad (\geq 0 \quad \checkmark)$$

and $E = \int d^{D-1} x \mathcal{H}$ is conserved.

The equations of motion are

$$\frac{\partial \mathcal{L}}{\partial \phi} - \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial \partial_\mu \phi} \right) = 0$$

$$(7.2) \quad \text{i.e.} \quad \partial^\mu \partial_\mu \phi = m^2 \phi$$

- This has plane-wave solutions

$$\phi(t, \vec{x}) = a_p e^{-iE_p t + i\vec{p} \cdot \vec{x}} + a_p^* e^{iE_p t - i\vec{p} \cdot \vec{x}} \quad (\phi \in \mathbb{C})$$

where

$$(7.3) \quad E_p = \sqrt{\vec{p}^2 + m^2}$$

- Fourier transforming is convenient:

$$\phi = \int \frac{d^D x}{(2\pi)^D} e^{i\vec{p} \cdot \vec{x}} \phi(\vec{p})$$

$$(7.4) \quad \Rightarrow \quad (p^2 + m^2) \phi(\vec{p}) = 0 \quad \forall \vec{p}$$

$p^2 + m^2 = 0$ is called "on-shell"
"on the mass shell".

Classically $\phi(\vec{p})$ vanishes off-shell.

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Let's now do it in light-cone coordinates.

Write $\vec{x}_T = (x^2, \dots, x^D)$ transverse coords
and

Fourier transform only in $\left\{ \begin{matrix} \text{transverse} \\ p^+ \end{matrix} \right\}$ directions

$$(8.1) \Rightarrow \phi(x^+, x^-, \vec{x}_T) = \int \frac{dp^+}{2\pi} \int \frac{d^{D-2} \vec{p}_T}{(2\pi)^{D-2}} e^{-ix^- p^+ + i\vec{x}_T \cdot \vec{p}_T} \phi(x^+, p^+, \vec{p}_T)$$

This satisfies

$$(8.2) \quad \left[i2\partial_+ - \frac{1}{2p^+} (m^2 + p^I p^I) \right] \phi(x^+, p^+, \vec{p}_T) = 0 \quad \text{first order}$$

Quantization

Let's write the field as a sum of plane-waves with operator coefficients this time :-

$$(8.3) \quad \boxed{\phi = \frac{1}{\sqrt{V}} \sum_{\vec{p}} \frac{1}{\sqrt{2E_p}} (a_p e^{-iE_p t + i\vec{p} \cdot \vec{x}} + a_p^\dagger e^{iE_p t - i\vec{p} \cdot \vec{x}})}$$

(8.4) In a box of volume $V = \prod_{i=1}^D L_i$, we require quantization $\boxed{p_i = \frac{2\pi n_i}{L_i}}$ (no sum)

Working through the algebra gives

$$S = \int dt \left(\dot{a}^*(t) \dot{a}(t) \cdot \frac{1}{2E_p} - \frac{1}{2} E_p a^*(t) a(t) \right)$$

$$(8.5) \quad \text{and } E = \int dt \left(\frac{1}{2E_p} \dot{a}^* \dot{a} + \frac{1}{2} E_p a^* a \right)$$

Canonical momenta

$$p = \frac{\partial \mathcal{L}}{\partial \dot{a}} = \dot{a}^* \quad ; \quad p^* = \frac{\partial \mathcal{L}}{\partial \dot{a}^*} = \dot{a}$$

(8.6) and they satisfy $\boxed{\dot{a} = -E_p^2 a}$

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This has the solution

$$a(t) = a_p e^{-iE_p t} + a_p^\dagger e^{iE_p t}$$

(9.1) So
$$H = E_p (a_p^\dagger a_p + a_{-p}^\dagger a_{-p})$$

Oscillators must satisfy canonical commutation relations

$$[a_p, a_p^\dagger] = 1 \quad ; \quad [a_p, a_{-p}^\dagger] = 1$$

These come about from replacing Poisson brackets with commutators. This is canonical quantization.

In general, we label our a 's with labels p that distinguish them as separate modes, so in fact

(9.2)
$$[a_p, a_k^\dagger] = \delta_{p,k} \quad ; \quad [a_p, a_k] = 0 = [a_p^\dagger, a_k^\dagger]$$

Modulo a zero-point energy (see later on!) we have

(9.3)
$$E = \sum_{\vec{p}} E_p a_p^\dagger a_p \quad , \quad \vec{p} = \sum_{\vec{p}} \vec{p} \underbrace{a_p^\dagger a_p}_{\text{number operator for } p\text{-th mode}}$$

What we have are simple harmonic oscillators.

Vacuum state (no particles): $|\Omega\rangle$

Build one-particle state with momentum \vec{p} by using $a_p^\dagger |\Omega\rangle$

This has momentum \vec{p} and energy E_p - by using (9.3).

D In light-cone gauge, one-particle states for this scalar field are specified by $(p^+, \vec{p}_\perp) \equiv p^- (p^+, \vec{p}_\perp)$.

(9.4)
$$\Rightarrow a_{p^+, \vec{p}_\perp}^\dagger |\Omega\rangle \Leftrightarrow \text{1-particle states}$$

Spin-one

(10.1) Relativistic notation combines \vec{E} and \vec{B} into $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$ and the dynamics is invariant under the gauge symmetry

(10.2) $A_\mu \rightarrow A_\mu + \partial_\mu \epsilon = A_\mu + \delta A_\mu$

(10.3) $S = \int d^4x \left(-\frac{1}{4} F^{\mu\nu} F_{\mu\nu} \right)$ (free field)

(10.4) whereupon $\partial_\mu F^{\mu\nu} = 0$

Using (10.1), this becomes

(10.5) $\partial^2 A^\nu - \partial^\nu (\partial \cdot A) = 0$

Again, Fourier transform:

(10.6) $A^\mu(x) = \int \frac{d^4p}{(2\pi)^4} e^{ip \cdot x} A(p)$

(10.7) then $p^2 A^\mu - p^\mu (p \cdot A) = 0$ (e-o-m)

Gauge transformation takes $\delta A_\mu(x) = \partial_\mu \epsilon$

So in momentum space this is

(10.8) $\delta A(p) = i p_\mu \epsilon(p)$

Reality of $\epsilon \Rightarrow \epsilon(-p) = \epsilon^*(p)$

In light-front gauge :-

Dynamical variables $\{A^+(p), A^-(p), A^\perp(p)\}$

(10.9) Pick $A_+(p) = 0$ (can always find an ϵ to do this, for this one component of A^μ).

Then e-o-m says

(10.10) $\oplus p^+ (A \cdot p) = 0$ i.e. $p \cdot A = 0 \Rightarrow A^- = \frac{1}{p^+} p^\perp A^\perp$

$\ominus p^2 A^\mu = 0$

Therefore, $A^\perp = 0$ off-shell. On-shell: $p^2 = 0$

(ii)

(11.1)

$$\text{One-photon states : } \sum_{I=2}^{D-1} \xi_I a_{p^+, p_T}^{I\dagger} |\Omega\rangle$$

The ξ^I is called the polarization vector.

Spin-two

Einstein's equations for GR, $G_{\mu\nu} = 0$, can be linearized, as can the general coordinate transformations.

(11.2)

The eqn of motion becomes ($g_{\mu\nu} \approx \eta_{\mu\nu} + h_{\mu\nu}$)

$$p^2 h^{\mu\nu} - p_\alpha (p^\mu h^{\nu\alpha} + p^\nu h^{\mu\alpha}) + p^\mu p^\nu h^\lambda{}_\lambda = 0$$

(11.3)

The gauge transformations are

$$\delta h^{\mu\nu}(p) = i p^\mu \epsilon^\nu(p) + i p^\nu \epsilon^\mu(p)$$

Light-cone gauge variables are $\{h^{++}, h^{+-}, h^{--}, h^{+I}, h^{-I}, h^{\pm J}\}$

By suitable gauge transformation, can set

$$h^{++} = h^{+-} = h^{+I} = 0 \quad (\text{like } A^+ = 0)$$

and the d.o.f. remaining are $(h^{--}, h^{-I}, h^{\pm J})$.

Find

$$h^I{}_I = 0 \quad (\text{traceless})$$

and

$$p^2 h^{\mu\nu} = 0$$

$$p_\alpha h^{\nu\alpha} = 0$$

\sim transverse

D.o.f. : symmetric traceless tensor

$$\# = \frac{1}{2} (D-2)(D-1) - 1 = \frac{1}{2} D(D-3)$$

Construct

$$\text{One-graviton states : } \sum_{I, J=2}^{D-1} \xi_{IJ} a_{p^+, p_T}^{IJ\dagger} |\Omega\rangle ; \xi_{II} = 0$$

ξ_{IJ} is symmetric traceless $\Rightarrow h(0)$ dof. on-shell.